

Geometric analysis of the Lorentzian distance function

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Geometric analysis of the Lorentzian distance function on trapped submanifolds, *Classical and Quantum Gravity* **33** (2016) 125007 (28 pp.).

- Our results were strongly based on a previous work by Erkekoglu, García-Río and Kupeli, where they established the basis for the comparison analysis of the (Lorentzian) Hessian and Laplacian operators of the Lorentzian distance function:

On level sets of Lorentzian distance function, *General Relativity and Gravitation* **35** (2003), 1597–1615.

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- Obviously, $p \ll q$ implies $p < q$. As usual, $p \leq q$ means that either $p < q$ or $p = q$.
- For a subset $S \subset M$, one defines the **chronological future of S** as

$$I^+(S) = \{q \in M : p \ll q \text{ for some } p \in S\},$$

and the **causal future of S** as

$$J^+(S) = \{q \in M : p \leq q \text{ for some } p \in S\}.$$

Thus $S \cup I^+(S) \subset J^+(S)$.

Causality relations

- In a dual way, $I^-(S) = \{q \in M : q \ll p \text{ for some } p \in S\}$ and $J^-(S) = \{q \in M : q \leq p \text{ for some } p \in S\}$ are the **chronological past** and **causal past** of S .

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- For instance, for a point $p \in \overline{M}$ in Minkowski space, $I^+(p)$ is just the future timecone of p ,

$$I^+(p) = \{q \in \overline{M} : \langle q - p, q - p \rangle < 0 \text{ and } \langle q - p, e_{n+1} \rangle < 0\},$$

and

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- $I^+(p)$ is always **open**. $J^+(p)$ is **neither open nor closed** in general.

Lorentzian distance function

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- Globally hyperbolic spacetimes turn out to be the natural class of spacetimes for which the Lorentzian distance function is **finite-valued and continuous**.
- Recall that a spacetime M is said to be **globally hyperbolic** if
 - (i) it is **causal**, that is, there exists no causal loop in M , and
 - (ii) the intersections $J^+(p) \cap J^-(q)$ are compact for every $p, q \in M$.

Lorentzian distance function from a point

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- Let $T_{-1}M|_p$ be the fiber of the **unit future observer bundle** of M at p , that is,

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- Define the function $s_p : T_{-1}M|_p \rightarrow [0, +\infty]$ by

$$s_p(v) = \sup\{t \geq 0 : d_p(\gamma_v(t)) = t\},$$

where $\gamma_v : [0, a) \rightarrow M$ is the future inextendible geodesic starting at p with initial velocity v .

Lorentzian distance function from a point

- Then, one can define the subset $\tilde{\mathcal{I}}^+(p) \subset T_p M$ given by

$$\tilde{\mathcal{I}}^+(p) = \{tv : \text{for all } v \in T_{-1}M|_p \text{ and } 0 < t < s_p(v)\}$$

and consider the subset $\mathcal{I}^+(p) \subset M$ given by

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Lemma 1 (Erkekoglu, García-Río and Kupeli, GRG 2003)

Let M be a spacetime and $p \in M$.

- 1 If M is **strongly causal** at p , then $s_p(v) > 0$ for all $v \in T_{-1}M|_p$ and $\mathcal{I}^+(p) \neq \emptyset$.

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- If M is **strongly causal** at p , then $s_p(v) > 0$ for all $v \in T_{-1}M|_p$ and $\mathcal{I}^+(p) \neq \emptyset$.
- If $\mathcal{I}^+(p) \neq \emptyset$, then the Lorentzian distance function d_p is **smooth** on $\mathcal{I}^+(p)$ and its gradient $\bar{\nabla} d_p$ is a **past-directed timelike (geodesic) unit** vector field on $\mathcal{I}^+(p)$.

Hessian comparison results for the Lorentzian distance

- For every $c \in \mathbb{R}$, let us define

$$h_c(t) = \begin{cases} \frac{1}{\sqrt{c}} \sinh(\sqrt{c} t) & \text{if } c > 0 \text{ and } t > 0 \\ t & \text{if } c = 0 \text{ and } t > 0 \\ \frac{1}{\sqrt{-c}} \sin(\sqrt{-c} t) & \text{if } c < 0 \text{ and } 0 < t < \pi/\sqrt{-c}. \end{cases}$$

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- Observe that the index of a Jacobi field along a timelike geodesic in a Lorentzian space form of constant curvature c is given by

$$I_{\gamma_c}(J_c, J_c) = -\frac{h'_c(t)}{h_c(t)} \langle x, x \rangle.$$

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$$I_{\gamma_c}(J_c, J_c) = -\frac{h'_c(t)}{h_c(t)} \langle X, X \rangle.$$

- On the other hand, $\frac{h'_c(t)}{h_c(t)}$ is the future mean curvature of the level set

$$\Sigma_c(t) = \{q \in \mathcal{I}^+(p) : d_p(q) = t\} \subset M_c^n.$$

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Lemma 2

Let M be a spacetime such that $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M . Assume that there exists a point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and let $q \in \mathcal{I}^+(p)$ (with $d_p(q) < \pi/\sqrt{-c}$ when $c < 0$). Then for every spacelike vector $x \in T_q M$ orthogonal to $\bar{\nabla} d_p(q)$

$$\bar{\nabla}^2 d_p(x, x) \leq -\frac{h'_c}{h_c}(d_p(q))\langle x, x \rangle,$$

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where $\bar{\nabla}^2$ stands for the Hessian operator on M .

- The proof of Lemma 2 follows from the fact that

$$\bar{\nabla}^2 d_p(x, x) = I_\gamma(J, J)$$

where γ is the radial future directed unit timelike geodesic from p to q and J is the Jacobi field along γ with $J(0) = 0$ and $J(s) = x$, and is strongly based on the maximality of the index of Jacobi fields.

Hessian comparison results for the Lorentzian distance

- On the other hand, under the assumption that the sectional curvatures of the timelike planes of M are bounded **from above** by a constant c , we get the following result.

Lemma 3

Let M be a spacetime such that $K_M(\Pi) \leq c$ $c \in \mathbb{R}$, for all timelike planes in M . Assume that there exists a point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and let $q \in \mathcal{I}^+(p)$ (with $d_p(q) < \pi/\sqrt{-c}$ when $c < 0$). Then for every spacelike vector $x \in T_q M$ orthogonal to $\bar{\nabla} d_p(q)$ it holds that

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- The proof is similar to that of Lemma 2.

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- Our first objective is to compute the Hessian of u . To do that, observe that

$$\bar{\nabla} r = \nabla u + (\bar{\nabla} r)^\perp$$

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- By Gauss and Weingarten formulae we get

$$\bar{\nabla}_X \bar{\nabla} r = \nabla_X \nabla u - A_{(\bar{\nabla} r)^\perp} X + \text{II}(X, \nabla u) + \nabla_X^\perp (\bar{\nabla} r)^\perp,$$

for every tangent vector $X \in T\Sigma$, where II denotes the second fundamental form of the submanifold and, for every normal vector η , A_η denotes the Weingarten endomorphism with respect to η .

- It follows from here that

$$\nabla^2 u(X, Y) = \bar{\nabla}^2 r(X, Y) + \langle \text{II}(X, Y), \bar{\nabla} r \rangle$$

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- Tracing this expression, one gets that the Laplacian of u is given by

$$\Delta u = \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + m \langle \text{H}, \bar{\nabla} r \rangle,$$

where $\{E_1, \dots, E_m\}$ is a local orthonormal frame on Σ , and

$$\text{H} := \frac{1}{m} \text{tr}(\text{II}) = \frac{1}{m} \sum_{i=1}^m \text{II}(E_i, E_i)$$

defines the mean curvature vector field of the submanifold.

Case 1: Spacelike hypersurfaces contained in $\mathcal{I}^+(p)$

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- Therefore, the Laplacian of u becomes in this case

$$\begin{aligned}\Delta u &= \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + mH \langle N, \bar{\nabla}r \rangle \\ &= \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + mH \sqrt{1 + |\nabla u|^2}.\end{aligned}$$

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- Before stating our main results, we need to introduce some terminology.

The Omori-Yau maximum principle

- Following the terminology introduced by Pigola, Rigoli and Setti (2005), the **Omori-Yau maximum principle** is said to hold on an n -dimensional Riemannian manifold Σ if, for any smooth function $u \in \mathcal{C}^2(\Sigma)$ with $u^* = \sup_{\Sigma} u < +\infty$ there exists a sequence of points $\{p_k\}_{k \in \mathbb{N}}$ in Σ with the properties

$$(i) \ u(p_k) > u^* - \frac{1}{k}, \quad (ii) \ |\nabla u(p_k)| < \frac{1}{k}, \quad \text{and} \quad (iii) \ \Delta u(p_k) < \frac{1}{k}.$$

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- Equivalently, for any $u \in \mathcal{C}^2(\Sigma)$ with $u_* = \inf_{\Sigma} u > -\infty$ there exists a sequence of points $\{p_k\}_{k \in \mathbb{N}}$ in Σ satisfying

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- In this sense, the **classical** maximum principle given by Omori (1967) and Yau (1975) states that the Omori-Yau maximum principle holds on every complete Riemannian manifold with **Ricci curvature bounded from below**.

First applications: Hypersurfaces bounded by a level set of the Lorentzian distance. Case $K_M(\Pi) \geq c$

Theorem 1 (Alías, Hurtado, Palmer, TAMS 2010)

Let M^{m+1} be an $(m+1)$ -dimensional spacetime such that $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M . Let $p \in M$ be such that $\mathcal{I}^+(p) \neq \emptyset$, and let $\psi: \Sigma^m \rightarrow M^{m+1}$ be a spacelike hypersurface such that $\psi(\Sigma) \subset \mathcal{I}^+(p)$. If the Omori-Yau maximum principle holds on Σ (and $\inf_{\Sigma} u < \pi/\sqrt{-c}$ when $c < 0$), then its future mean curvature H satisfies

$$\sup_{\Sigma} H \geq f_c(\inf_{\Sigma} u),$$

where u denotes the Lorentzian distance d_p along the hypersurface. In particular, if $\inf_{\Sigma} u = 0$ then $\sup_{\Sigma} H = +\infty$.

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Corollary 1 (Alías, Hurtado, Palmer, TAMS 2010)

Under the assumptions of Theorem 1, if H is bounded from above on Σ , then there exists some $\delta > 0$ such that $\psi(\Sigma) \subset O^+(p, \delta)$, where $O^+(p, \delta)$ denotes the **future outer ball** of radius δ ,

$$O^+(p, \delta) = \{q \in I^+(p) : d_p(q) > \delta\}.$$

Proof of Theorem 1

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- The last assertion follows from the fact that $\lim_{s \rightarrow 0} f_c(s) = +\infty$.

Hypersurfaces in Lorentzian space forms

Theorem 2 (Alías, Hurtado, Palmer, TAMS 2010)

Let M_c^{m+1} be a Lorentzian space form of **constant sectional curvature c** and let $p \in M_c^{m+1}$. Let us consider $\psi : \Sigma^m \rightarrow M_c^{m+1}$ a spacelike hypersurface such that $\psi(\Sigma) \subset \mathcal{I}^+(p) \cap B^+(p, \delta)$ for some $\delta > 0$ (with $\delta \leq \pi/\sqrt{-c}$ if $c < 0$). If the Omori-Yau maximum principle holds on Σ , then

$$\inf_{\Sigma} H \leq f_c(\sup_{\Sigma} u) \leq f_c(\inf_{\Sigma} u) \leq \sup_{\Sigma} H,$$

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Corollary 2 (Alías, Hurtado, Palmer, TAMS 2010)

Let M_c^{m+1} be a Lorentzian space form of **constant sectional curvature c** and let $p \in M_c^{m+1}$. If Σ is a **complete** spacelike hypersurface in M_c^{m+1} with **constant mean curvature H** which is contained in $\mathcal{I}^+(p)$ and bounded from above by a level set of the Lorentzian distance function d_p (with $d_p < \pi/\sqrt{-c}$ if $c < 0$), then Σ is necessarily a level set of d_p .

- For a proof simply observe that the Ricci curvature of a spacelike hypersurface in an arbitrary spacetime M is given by

$$\begin{aligned} \text{Ric}(X, X) &= \text{Ric}_M(X, X) - \left(K_M(X \wedge N) + \frac{m^2 H^2}{4} \right) |X|^2 + |AX + \frac{m}{2} X| \\ &\geq \text{Ric}_M(X, X) - \left(K_M(X \wedge N) + \frac{m^2 H^2}{4} \right) |X|^2. \end{aligned}$$

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Corollary 3 (Alías, Hurtado, Palmer, TAMS 2010)

The only **complete** spacelike hypersurfaces with **constant mean curvature** in the Lorentz-Minkowski space \mathbb{L}^{m+1} which are contained in $\mathcal{I}^+(p)$ (for some fixed $p \in \mathbb{L}^{m+1}$) and bounded from above by a hyperbolic space centered at p are precisely the hyperbolic spaces centered at p .

Hyperbolicity of spacelike hypersurfaces

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- (i) If the future mean curvature of Σ satisfies $H \leq \frac{2\sqrt{m-1}}{m} f_c(u)$ with $H < f_c(u)$ at some point of Σ if $m = 2$, then Σ is hyperbolic.
- (ii) If $c = 0$ and $H \leq 0$, then Σ is hyperbolic.
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In particular, every **maximal** hypersurface contained in $\mathcal{I}^+(p)$ (and satisfying $u < (\pi/2\sqrt{-c})$ if $c < 0$) is hyperbolic.

Proof of Theorem 3

- In order to proof (i), observe that under our assumptions on H we have

$$mH \leq 2\sqrt{m-1} f_c(u) \leq \frac{f_c(u)(m + |\nabla u|^2)}{\sqrt{1 + |\nabla u|^2}}.$$

- In order to proof (i), observe that under our assumptions on H we have

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- Therefore, by the Hessian comparison result in Lemma 2 we conclude

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- To prove (ii) and (iii), simply observe that $f_0(u) = 1/u > 0$ and $f_c(u) = \sqrt{c} \coth(\sqrt{c}u) > \sqrt{c}$ on Σ .

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$$\phi_c(t) = \begin{cases} \frac{1}{c} \cosh(\sqrt{c} t) & \text{if } c > 0 \text{ and } t > 0 \\ \frac{t^2}{2} & \text{if } c = 0 \text{ and } t > 0 \\ \frac{1}{c} \cos(\sqrt{-c} t) & \text{if } c < 0 \text{ and } 0 < t < \pi/\sqrt{-c}. \end{cases}$$

- Then, the Laplacian of v is given by

$$\begin{aligned} \Delta v &= \phi'_c(u) \Delta u + \phi''_c(u) |\nabla u|^2 \\ &= h_c(u) \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + m h_c(u) \langle H, \bar{\nabla} r \rangle + h'_c(u) |\nabla u|^2. \end{aligned}$$

Case 2: Spacelike submanifolds contained in $\mathcal{I}^+(p)$

- Recall that the Laplacian of u is given by

$$\Delta u = \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + m \langle H, \bar{\nabla} r \rangle.$$

- Consider the function $v = \phi_c(u)$, where $\phi_c(t)$ is a primitive of $h_c(t)$:

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- Assume now that $K_M(\Pi) \geq c$ (resp. $K_M(\Pi) \leq c$) for all timelike planes in M .

- Then by the Hessian comparison results for r , one gets that

$$\overline{\nabla}^2 r(X, X) \leq (\geq) - \frac{h'_c}{h_c}(u)(1 + \langle X, \nabla u \rangle^2)$$

for every unit tangent vector field $X \in T\Sigma$.

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- Therefore,

$$h_c(u) \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) \leq (\geq) - h'_c(u)(m + |\nabla u|^2),$$

which, jointly with the expression above, gives the following inequality for the Laplacian of v

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- Summarizing:

- $K_M(\Pi) \geq c$ implies that

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where $v = \phi_c(u)$ and u is the Lorentzian distance function of M restricted on the spacelike submanifold Σ .

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- For statement of our main results, we introduce some terminology. 



The weak maximum principle

- The *weak maximum principle* is said to hold on Σ if, for any $u \in \mathcal{C}^2(\Sigma)$ with $u^* < +\infty$ there is a sequence $\{p_k\}_{k \in \mathbb{N}}$ in Σ with

$$(i) \quad u(p_k) > u^* - \frac{1}{k}, \quad \text{and} \quad (iii) \quad \Delta u(p_k) < \frac{1}{k}.$$

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- This is equivalent (among other conditions) to the fact that for every $\lambda > 0$, the only non-negative bounded smooth solution u of $\Delta u \geq \lambda u$ on Σ is the constant $u = 0$.
- In particular, every **parabolic** manifold is stochastically complete. Hence, the weak max principle holds on every parabolic manifold.

Trapped submanifolds in a spacetime

- Following the standard terminology in General Relativity, a spacelike submanifold Σ^m (of arbitrary codimension) of a spacetime M^n is said to be a **future trapped** submanifold if its mean curvature vector field H is **timelike and future-pointing** everywhere on Σ .

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- On the other hand, if H is **lightlike and future-pointing** everywhere on Σ then the spacelike submanifold is said to be **marginally future trapped**.

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- Finally, Σ is said to be **weakly future trapped** if H is **causal** (that is, timelike or lightlike) and **future-pointing** everywhere.

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- Finally, Σ is said to be **weakly future trapped** if H is **causal** (that is, timelike or lightlike) and **future-pointing** everywhere.
- Analogously, Σ is said to be **weakly past trapped** if H is **causal and past-pointing** on Σ .

Weakly trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

Theorem 4 (Alías, Bessa, de Lira, CQG 2016)

Let M be a spacetime with a reference point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and assume $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M .

- 1 If $c \geq 0$ there exists no stochastically complete, weakly past trapped submanifold contained in $\mathcal{I}^+(p)$.
- 2 If $c < 0$ and Σ is a stochastically complete, weakly past trapped submanifold contained in $\mathcal{I}^+(p) \cap B^+(p, \pi/\sqrt{-c})$, then

$$u_* = \inf_{\Sigma} u \geq \frac{\pi}{2\sqrt{-c}},$$

where u denotes the Lorentzian distance d_p along the hypersurface. In other words, Σ is contained in $B^+(p, \pi/\sqrt{-c}) \cap O^+(p, \pi/2\sqrt{-c})$.

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- Recall that the subsets $B^+(p, \delta)$ and $O^+(p, \delta)$ denote the **future inner ball** and the **future outer ball** of radius $\delta > 0$, that is,

$$B^+(p, \delta) = \{q \in I^+(p) : d_p(q) < \delta\}$$

$$O^+(p, \delta) = \{q \in I^+(p) : d_p(q) > \delta\}.$$

Proof of Theorem 4

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$$-\frac{1}{k} < \Delta v(p_k) \leq -mh'(u(p_k)) + mh(u(p_k))\langle H, \bar{\nabla} r \rangle(p_k),$$

for $\{p_k\} \subset \Sigma$ with $\lim_{k \rightarrow \infty} v(p_k) = v_*$ and $\lim_{k \rightarrow \infty} u(p_k) = u_*$.

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and, making $k \rightarrow \infty$ here we get $h'_c(u_*) \leq 0$.

- The result then follows by observing that, when $c \geq 0$ then $h'_c(t) > 0$, and if $c < 0$ then $h'_c(t) \leq 0$ when $\pi/2\sqrt{-c} \leq t < \pi/\sqrt{-c}$.

Marginally trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

Theorem 5 (Alías, Bessa, de Lira, CQG 2016)

Let M be a spacetime with a reference point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and assume $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M . Let Σ be a stochastically complete, marginally trapped submanifold contained in $\mathcal{I}^+(p)$ (with $u_* < \pi/2\sqrt{-c}$ in the case $c < 0$). Then

$$\sup_{\Sigma} |H_0| \geq \frac{h'_c}{h_c}(u_*),$$

where H_0 stands for the spacelike component of the lightlike vector field H which is orthogonal to $\bar{\nabla}r$, and $u_* = \inf_{\Sigma} u$. In particular, if $u_* = 0$ then $\sup_M |H_0| = +\infty$.

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Corollary 4 (Alías, Bessa, de Lira, CQG 2016)

Under the assumptions of Theorem 5, if $|H_0|$ is bounded from above on Σ , then there exists some $\delta > 0$ such that $\Sigma \subset O^+(p, \delta)$, where $O^+(p, \delta)$ denotes the **future outer ball** of radius δ .

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$$\langle H, \bar{\nabla} r \rangle = |H_0| > 0 \quad \text{on } \Sigma.$$

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- Otherwise, let us write

$$\Delta v \leq -mh'_c(u) + mh_c(u)|H_0| \leq -mh'_c(u) + mh_c(u) \sup_{\Sigma} |H_0|.$$

Proof of Theorem 5

- Applying again the weak maximum principle on Σ to the function $v = \phi_c(u)$, with $v_* = \inf_{\Sigma} v = \phi_c(u_*)$, we have

$$-\frac{1}{k} < \Delta v(p_k) \leq -mh'_c(u(p_k)) + mh_c(u(p_k)) \sup_{\Sigma} |H_0|,$$

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$$\sup_{\Sigma} |H_0| \geq \frac{h'_c(u_*)}{h_c(u_*)}.$$

- The last assertion follows from the fact that $h_c(0) = 0$ and $h'_c(0) = 1$.

Marginally trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \leq c$

Theorem 6 (Alías, Bessa, de Lira, CQG 2016)

Let M be a spacetime with a reference point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and assume $K_M(\Pi) \leq c$, $c \in \mathbb{R}$, for all timelike planes in M . Let Σ be a stochastically complete, marginally future trapped submanifold contained in $\mathcal{I}^+(p) \cap B^+(p, \delta)$ for some $\delta > 0$ (with $\delta \leq \pi/\sqrt{-c}$ when $c < 0$). Then

$$\inf_{\Sigma} |H_0| \leq \frac{h'_c}{h_c}(u^*),$$

where H_0 stands for the spacelike component of the lightlike vector field H which is orthogonal to $\overline{\nabla}r$, and $u^* = \sup_{\Sigma} u$.

Proof of Theorem 6

- Since $K_M(\Pi) \leq c$ and $\langle H, \bar{\nabla} r \rangle = |H_0| > 0$ on Σ , we have

$$\Delta v \geq -mh'_c(u) + mh_c(u)|H_0|.$$

Proof of Theorem 6

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$$\Delta v \geq -mh'_c(u) + mh_c(u)|H_0|.$$

- If $\inf_{\Sigma} |H_0| = -\infty$ then there is nothing to prove.

Proof of Theorem 6

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for $\{p_k\} \subset \Sigma$ with $\lim_{k \rightarrow \infty} v(p_k) = v^*$ and $\lim_{k \rightarrow \infty} u(p_k) = u^*$.

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- Making $k \rightarrow +\infty$ we conclude that

$$\inf_{\Sigma} |H_0| \leq \frac{h'_c(u^*)}{h_c(u^*)}.$$

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Let M_c^n be a Lorentzian space form of **constant sectional curvature c** and let $p \in M_c^n$. Let Σ be a stochastically complete, marginally trapped submanifold of M_c^n which is contained in $\mathcal{I}^+(p) \cap B^+(p, \delta)$ for some $\delta > 0$ (with $\delta \leq \pi/2\sqrt{-c}$ if $c < 0$). Then

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- The estimates are sharp as proved by considering Σ as a constant mean curvature hypersurface of a level set of the Lorentzian distance in M_c^n .

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and congratulations to Eduardo for his forthcoming first 60 years.